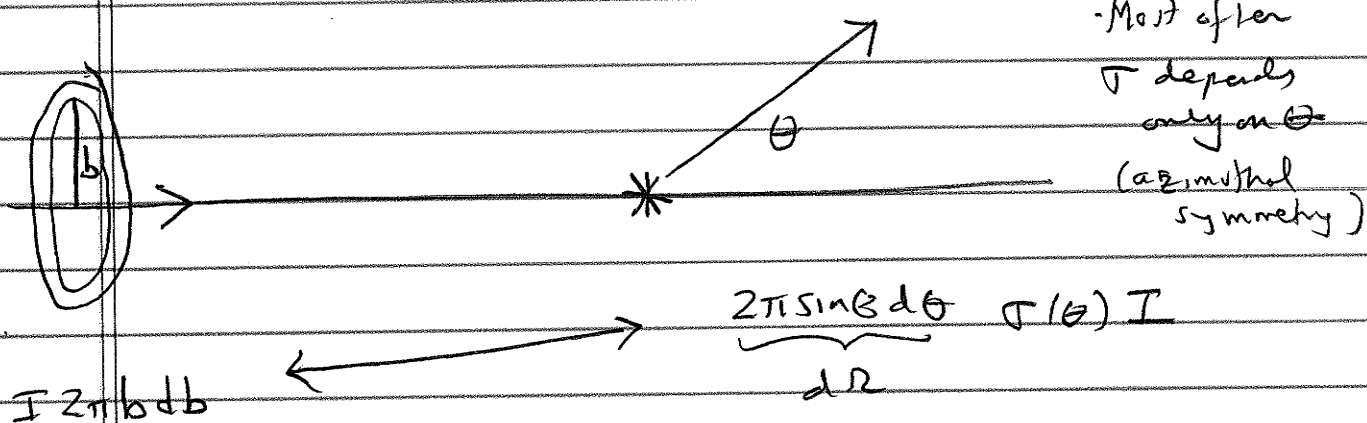


Recall scattering theory definitions

$$d\sigma = \sigma(\Omega) d\Omega \equiv \frac{\# \text{ particles scattered into solid angle } d\Omega \text{ per unit time}}{\text{incident intensity}}$$

In terms of "impact parameter"  $b$



- Most often  
 $\sigma$  depends  
 only on  $\theta$   
 (azimuthal  
 symmetry)

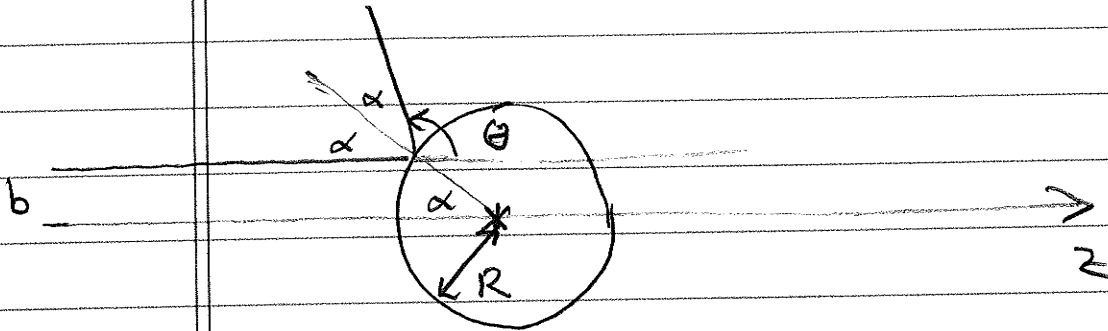
$$\sigma(\theta) = \frac{b}{\sin\theta} \left| \frac{db}{d\theta} \right|$$

We did one problem: very fast incoming particles  
 going past  $1/r$  potential.

"Total cross section"

$$\sigma = \int \sigma(\theta) d\Omega$$

Another famous scattering problem is hard sphere



$$\theta = 0 \text{ if } b > R$$

$$\sin \alpha = b/R \quad \theta = \pi - 2\alpha \quad \alpha = \frac{\pi}{2} - \frac{\theta}{2}$$

$$b = R \sin \alpha = R \sin\left(\frac{\pi}{2} - \frac{\theta}{2}\right)$$

$$b = R \cos \theta/2 \quad \text{check limits: } b=0 \quad \theta=\pi$$

$$b=R \quad \theta=0$$

✓

$$\sigma(\theta) = \frac{b}{\sin \theta} \frac{db}{d\theta}$$

$$= \frac{R \cos \theta/2}{\sin \theta} \left| -R \frac{1}{2} \sin \theta/2 \right|$$

$$= \frac{1}{4} R^2 \quad \leftarrow \text{indep of } \theta \text{ (not typical)}$$

=

$$\sigma = \int \sigma(\theta) d\Omega = \int \frac{1}{4} R^2 2\pi \sin \theta d\theta$$

$$= \frac{1}{4} R^2 4\pi = \pi R^2 \quad \checkmark$$

SC-1

In QM imagine plane wave incoming

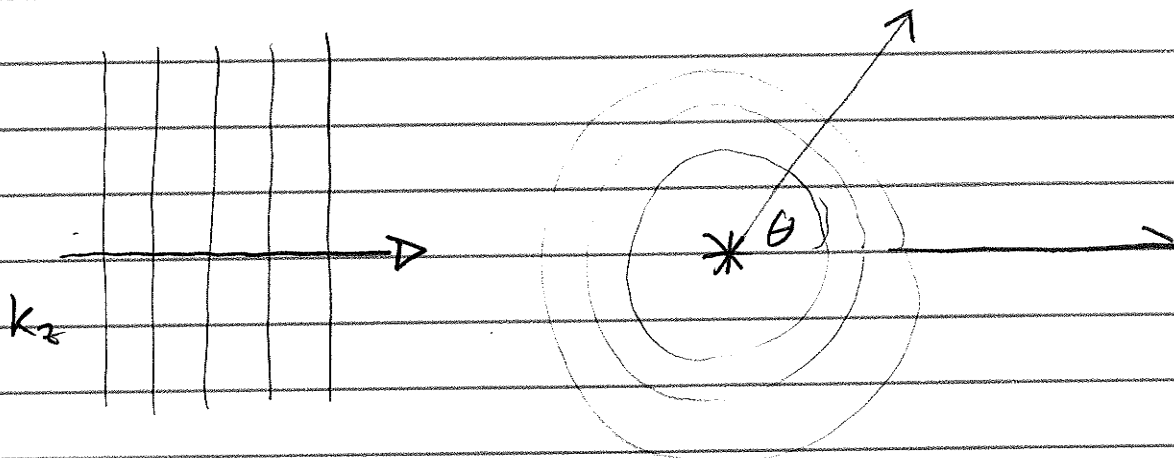
instead of particle (because plane wave is soln of

Sch: eqn) then hits scattering center. Look for

solution of form

$$\psi(\vec{r}) = A \left\{ e^{ikz} + f(\theta) \frac{e^{ikr}}{r} \right\}$$

"scattering amplitude"



$$k = \sqrt{2mE} / \hbar$$

probability density

Incident flux is  $|A|^2 v dt d\sigma$

(usual  $I = nevA$  argument)

Scattered flux is  $\frac{|A|^2 |f|^2}{r^2} v dt r^2 d\Omega$

Setting these equal  $\sigma(\theta) = \frac{d\sigma}{d\Omega} = |f|^2$

Sc-2

Schrodinger Eqn in spherical potential

$$[\hat{H}, \hat{L}^2] = [\hat{H}, \hat{L}_z] = 0$$

$$\psi(r, \theta, \phi) = R(r) Y_{lm}(\theta, \phi)$$

$$-\frac{\hbar^2}{2m} \frac{d^2}{dr^2} u + \left\{ V(r) + \frac{\hbar^2 l(l+1)}{2m r^2} \right\} u = E u$$

where  $u(r) = R(r) r$

Very large  $r$   $V(r) \rightarrow 0$  as does  $\frac{\hbar^2 l(l+1)}{2m r^2}$

$$\frac{d^2 u}{dr^2} = -k^2 u$$

$$u(r) = C e^{ikr} + D e^{-ikr}$$

↑  
outgoing

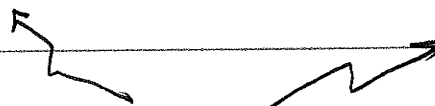
$$R(r) = e^{ikr}/r$$

further motivating our guess of form for solution

Assume  $V(r)$  is localized, i.e. there is a region

where  $V(r)$  might be big, but also an intermediate

region where  $V(r)$  can be ignored but  $\frac{\hbar^2 l(l+1)}{2m r^2}$  cannot



Might seem problematic  
eg Coulomb  $1/r$  but  
mostly screened Yukawa  $e^{-dr}/r$

Q: Does anyone recall solns to Radial Egn for  $V=0$ ?

A:  $R(r) = A j_l(kr) + B n_l(kr)$

↙ ↘  
Bessel functions.

We emphasized Bessel a lot like  $\sin, \cos$   
except they decay and have unevenly spaced roots.

Another similarity  $\sin \theta, \cos \theta \leftrightarrow e^{\pm i\theta}$   
 $e^{i\theta} = \sin \theta + i \cos \theta$

"Hankel functions"  $h_l^1 \equiv j_l + i n_l$

$$h_l^2 \equiv j_l - i n_l$$

$$h_l^1 \sim e^{ikr}/r \quad h_l^2 \sim e^{-ikr}/r \quad (i)$$

$$\psi(r, \theta, \phi) = A \left[ e^{ikz} + \sum_{lm} C_{lm} h_l^1(kr) Y_{lm}(\theta, \phi) \right]$$

Analogy: particle in a box  $-\frac{\hbar^2}{2m} \frac{d^2\psi}{dx^2} = E\psi$



$$\psi(x) = e^{\pm ikx}$$

$$E = \frac{\hbar^2 k^2}{2m}$$

But choose  $\sin kx$  / linear combo to satisfy boundary conditions

SC-3A

## Spherical Hankel functions

$$h_0^1 = -i \frac{e^{ix}}{x}$$

$$h_1^1 = \left( -\frac{i}{x^2} - \frac{1}{x} \right) e^{ix}$$

$$h_2^1 = \left( -\frac{3i}{x^2} - \frac{3}{x^2} + \frac{i}{x} \right) e^{ix}$$

$$h_0^2 = i \frac{e^{-ix}}{x}$$

$$h_1^2 = \left( \frac{i}{x^2} - \frac{1}{x} \right) e^{-ix}$$

$$h_2^2 = \left( \frac{3i}{x^2} - \frac{3}{x^2} + \frac{i}{x} \right) e^{-ix}$$

$$j_0 = \frac{\sin x}{x}$$

$$y_0 = -\frac{\cos x}{x}$$

$$j_1 = \frac{\sin x}{x^2} - \frac{\cos x}{x}$$

$$y_1 = -\frac{\cos x}{x^2} - \frac{\sin x}{x}$$

etc

$$j_l \sim x^l \frac{2^l l!}{(2l+1)!}$$

$$y_l \sim \frac{1}{x^{l+1}} \left( \frac{-2^l l!}{2^l l!} \right)$$

x small

Sc-4

$$Y_{\ell m}(\theta, \phi) \sim P_{\ell}(\cos \theta) \text{ for } m=0$$

Spherically symmetric  $\rightarrow$  no  $\phi$  dependence

$$\psi(r, \theta) = A \left[ e^{ikz} + k \sum_{\ell=0}^{\infty} i^{\ell+1} (2\ell+1) a_{\ell} h'_{\ell}(kr) P_{\ell}(\cos \theta) \right]$$

We said  $h'_{\ell}(kr) \sim \frac{e^{ikr}}{r}$  larger  $r$

These various factors

In fact  $h'_{\ell}(kr) = \frac{e^{ikr}}{kr} (-i)^{\ell+1}$  larger  $r$

Just a question of normalization of  $a_{\ell}$ .

We will see why this

is convenient

Very large  $r$

$$\psi(r, \theta) = A \left[ e^{ikz} + f(\theta) \frac{e^{ikr}}{r} \right]$$

$$f(\theta) = \sum_{\ell=0}^{\infty} (2\ell+1) a_{\ell} P_{\ell}(\cos \theta)$$

"partial wave amplitude"

Total cross section

$$\sigma = \int |f|^2 d\Omega$$

$$= \int \sum_{\ell} \sum_{\ell'} (2\ell+1)(2\ell'+1) a_{\ell} a_{\ell'} P_{\ell}(\cos \theta) P_{\ell'}(\cos \theta) d\Omega$$

$$= 4\pi \sum_{\ell=0}^{\infty} (2\ell+1) |a_{\ell}|^2$$

$$\int_{-1}^1 P_{\ell}(x) P_{\ell'}(x) dx = \frac{2}{2\ell+1} \delta_{\ell\ell'}$$

It is better to write everything including

Incoming wave in same form. Recall

$$e^{ikz} = \sum_{l=0}^{\infty} i^l (2l+1) j_l(kr) P_l(\cos\theta)$$

$$\psi(r, \theta) = A \sum_{l=0}^{\infty} i^l (2l+1) [j_l(kr) + ik a_l h_l'(kr)] P_l(\cos\theta)$$

Strategy for scattering problems:

① Solve Schrodinger eqn where  $V(r) \neq 0$   
to determine coefficients  $a_l$

②  $a_l \rightarrow f(\theta) \rightarrow \sigma(\theta)$

$a_l \rightarrow \sigma$

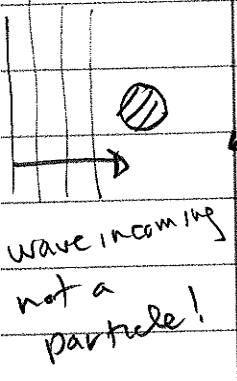


5C-6

Ex: Hard sphere scattering

HW 5-4 is finite depth problem.

$$V(r) = \begin{cases} \infty & r \leq a \\ 0 & r > a \end{cases}$$



Q: Any guesses concerning total  $\sigma$ ?

$$\psi(a, \theta) = 0$$

$$\sum_{l=0}^{\infty} i^l (2l+1) [j_l(ka) + ik a_e h_l'(ka)] P_l(\cos\theta) = 0$$

Since  $P_l$  are indep, the individual coefficients must vanish

$$a_e = i j_l(ka) / k h_l'(ka)$$

$$\sigma = \frac{4\pi}{k^2} \sum_{l=0}^{\infty} (2l+1) \left| \frac{j_l(ka)}{h_l'(ka)} \right|^2$$

Not much intuition here because who understands

$j_l$  and  $h_l'$  ?!?!

Consider low energy scattering  $ka \ll 1$  (like 5-4) HW

$$\frac{j_l(z)}{h_l'(z)} = \frac{j_l(z)}{j_l(z) + i\eta_l(z)} \approx -i \frac{j_l(z)}{\eta_l(z)}$$

$\uparrow$   
 $z$  small

sc-7

$$\frac{j_e(ka)}{h_e'(ka)} \sim \frac{i}{2l+1} \left[ \frac{2^l l!}{(2l)!} \right]^2 (ka)^{2l+1}$$

$$\sigma = \frac{4\pi}{k^2} \sum_{l=0}^{\infty} \frac{1}{(2l+1)} \left[ \frac{2^l l!}{(2l)!} \right]^4 (ka)^{4l+2}$$

$ka \ll 1$  keep only  $l=0$

$$\sigma = 4\pi a^2$$

↳

four times classical answer

"diffraction" of wave,

SC-18

"Born Approximation" is an alternate approach to scattering theory which is perturbative in  $V$ . Starting point is to recast Sch Eqn as integral eqn.

Another Example of diff eqn  $\leftrightarrow$  integral eqn

is provided by the starting point of scattering theory

Consider - Sch. Eqn in position representation

$$-\frac{\hbar^2}{2m} \nabla^2 \psi(r) + V(r) \psi(r) = E \psi(r)$$

Work instead in momentum space

$$\psi(r) = \frac{1}{(2\pi)^{3/2}} \int \psi(k_1) e^{-ik_1 r} d^3k_1$$

$$\frac{1}{(2\pi)^{3/2}} \int d^3k_1 \frac{\hbar^2 k_1^2}{2m} \psi(k_1) e^{-ik_1 r} + V(r) \left( \frac{d^3k_1}{(2\pi)^{3/2}} e^{-ik_1 r} \psi(k_1) \right)$$

$$= E \int \frac{d^3k_1}{(2\pi)^{3/2}} e^{-ik_1 r} \psi(k_1)$$

Multiply by  $e^{ikr}$  and  $\int d^3r / (2\pi)^3$

Using identity  $\int \frac{d^3r}{(2\pi)^3} e^{i(k-k_1)r} = \delta(k-k_1)$

SC-9

$$\frac{\hbar^2 k^2}{2m} \psi(k) + \frac{1}{(2\pi)^3} \int d^3 r V(r) e^{i k r} \int d^3 k_1 e^{-i k_1 r} \psi(k_1) = E \psi(k)$$

$$\int d^3 k_1 \frac{1}{(2\pi)^3} \int d^3 r e^{i(k-k_1)r} V(r) \psi(k_1)$$

$$\equiv V(k-k_1)$$

$$\rightarrow \frac{\hbar^2 k^2}{2m} \psi(k) + \int d^3 k_1 V(k-k_1) \psi(k_1) = E \psi(k)$$

E(k)

A physically pleasing eqn  $\psi(k_1)$  gets linked

to  $\psi(k)$  by Fourier component of potential V

at  $k-k_1$ !

$$(E(k) - E) \psi(k) = - \int d^3 k_1 V(k-k_1) \psi(k_1)$$

Can imagine trying to solve this in a similar

way to our derivation of time evolution operator...

SC-10

Although this integral Eqn for  $\psi(k)$  is enlightening, it is after most simple to work

in real space. The analogous formula is

$$\psi(\vec{r}) = \psi_0(\vec{r}) - \frac{m}{2\pi\hbar^2} \int \frac{e^{ik|\vec{r}-\vec{r}_0|}}{|\vec{r}-\vec{r}_0|} V(\vec{r}_0) \psi(\vec{r}_0) d^3r_0$$

soln to free  
particle Sch. Eqn  
( $V=0$ )

In words: soln for  $\psi$  at  $\vec{r}$  depends on value at  $\vec{r}_0$  times a "propagator" from  $\vec{r}_0$  to  $\vec{r}$ , and integrated over all space  $d^3r_0$

very  
This formula is useful because it allows an

iterative (perturbative) soln.

$$\psi(r) = A e^{ikz} - \frac{m}{2\pi\hbar^2} \frac{e^{ikr}}{r} \int e^{-i\vec{k}\cdot\vec{r}_0} V(\vec{r}_0) \psi(\vec{r}_0) d^3r_0$$

using  $\frac{1}{|\vec{r}-\vec{r}_0|} \sim \frac{1}{r}$  and  $e^{i\vec{k}|\vec{r}-\vec{r}_0|} \approx e^{ikr} e^{-i\vec{k}\cdot\vec{r}_0}$

Sc-11

$$\Rightarrow f(\theta, \phi) = -\frac{m}{2\pi\hbar^2 A} \int e^{-i\vec{k} \cdot \vec{r}_0} V(r_0) \psi(r_0) d^3 r_0$$

Born approx replaces  
this by  $A e^{i\vec{k} \cdot \vec{r}_0}$

$$= A e^{i\vec{k}' \cdot \vec{r}_0}$$

with  $\vec{k}' = k \hat{z}$

$$f(\theta, \phi) = -\frac{m}{2\pi\hbar^2} \int e^{i(\vec{k}' - \vec{k}) \cdot \vec{r}_0} V(r_0) d^3 r_0$$